Electromagnetic Waves

THE ELECTROMAGNETIC SPECTRUM						
Name	Frequency Range (Hz)	Wavelength Range	Production	Detection	Main Properties and uses	
Radiowaves	10 ⁴ to 10 ⁸	> 0 .1 m	Rapid acceleration and decelerations of electrons in aerials/ antenna.	Receiver's aerials.	Different specialised uses in radio communication.	
Microwaves	10^9 to 10^{12}	0.1 m te lmm	Klystron valve or magnetron valve	Point contact diodes.	 (a) Radar communication. (b) Analysis of fine details of molecular and atomic structure. 	
Infrared waves	3×10^{11} to 4×10^{14}	1mm te 700nm	Vibration of atoms and molecules.	Thermopiles Bolometer, Infra- red and photo- graphic film.	 (a) Useful for molecular structure. (b) Useful for haze photography. 	
Visible rays	4×10^{14}	700 nm to 400 nm	Electrons in atoms emit light when they move from one energy level to a lower energy level.	Human eye, Photocells and Photographic film.	Our eyes are sensitive to this range of wavelengths.	
Ultravielet rays	to 7 × 1 \bullet ¹⁴ 10 ¹⁴ to 10 ¹⁶	400 nm to 1 nm	Inner shell electrons in atoms moving from one energy level to a lower level.	Photocells and Photographic film.	 (a) Absorbed by glass (b) Can cause the tanning of the human skin. (c) Ionize atoms in atmo- sphere, resulting in the ionosphere. 	
Х-гауз	10 ¹⁶ to 10 ¹⁹	1 nm to 10 ⁻³ nm	X-ray tubes or inner shell electrons.	Photographic film, Geiger tubes and Ionization chamber.	 (a) Penetrate matter (e.g., radiography) (b) Ionize gases (c) Cause fluorescence (d) Cause photoelectric emission from metals. 	
Gamma rays	1• ¹⁸ to 10 ²²	< 1() ⁻³ nm	Radioactive decay of the nucleus.	Photographic film, Geiger tubes and Ionization chamber.	Similar to X-rays.	

Illustration 4

Obtain the photon energy in units of eV for different parts of the Em spectrum. In what way are the different scales of photon energies that you obtain related to the sources of electromagnetic radiation?

Sola.: The photon of given energy are released during transition between energy levels in the atom and the emitted photon energy is equal to difference of energies of energy levels, among which the transition has taken place. For $\lambda = 1$ m

$$E = hv = \frac{hc}{\lambda} = \frac{6.63 \times 10^{-34} \times 3 \times 10^8}{1} \text{ J}$$
$$E = \frac{19.89 \times 10^{-26}}{1.6 \times 10^{-19}} \text{ eV} = 12.425 \times 10^{-7} \text{ eV}$$

So, for emission of radio waves the energy difference

between energy levels should be $12.425 \times 10^{-7} \text{ eV}$.

Similarly, photon energy for other wavelengths can also be calculated

Type of Radiation		Wavelength of Radiation	Photon Energy
1.	Radio waves	1 m	$12.425 \times 10^{-7} eV$
2.	Microwaves	$3 \times 10^{-3} \mathrm{m}$	4.14× 10 ⁻⁴ eV
3.	Infrared waves	3×10 ⁻⁵ m	4.14 × 1 0 ⁻² eV
4.	Visible rays	$5 \times 10^{-7} \mathrm{m}$	2.485 eV
5.	Ultraviolet rays	$3 \times 10^{-7} \mathrm{m}$	4.14eV
6.	X-rays	10 ⁻¹⁰ m	$12.425 \times 10^3 \text{eV}$
7.	γ-rays	$10^{-12}{\rm m}$	$12.425 \times 10^{5} \text{eV}$

So, for emission of γ rays the energy difference among energy levels should be of the order of MeV, where for visible radiation it should be of the order of eV.