$n=2$ for second bright fringe and so on
$d=$ distance between two slits
$D=$ distance of sliw from the screen
$x_{n}=$ dismance of $n^{\text {th }}$ bright finge from the centre.
$\therefore x_{n}=\hat{n} \frac{D}{2}$

- For destracsive interference (i.e. formation of dark fringes).
For $n^{\text {ti }}$ dark fringe,
Path difierence $=x_{n} \frac{\frac{a}{D}}{D}=(2 n-1) \frac{\lambda}{2}$
where
$n=1$ for first dark ringe,
$n=2$ for $2^{\text {mi }}$ dark fringe and so on.
$x_{n}=$ distance of $n^{\text {it }}$ dark fringe from the centre
$\therefore x_{n}=(2 n-1) \frac{\lambda}{2} \frac{D}{d}$
- Fringe winth : The distance between any two consecutive bright or dark fringes is known as fringe width.
Fringe with , $\beta=\frac{\lambda D}{d}$
- Angular fringe width, $\theta=\frac{\beta}{D}=\frac{\lambda}{d}$
- If $\mathbb{F}_{1}, W_{2}$ are widths of two slits, $I_{1}, I_{2}$ are intensities of light coming from two slits; $a, b$ are the amplitudes of light from these slits, then.
$\frac{W_{1}}{W_{2}}=\frac{I_{1}}{I_{2}}=\frac{a^{2}}{b^{2}}$
$\frac{I_{\max }}{I_{\min }}=\frac{(a+\dot{b})^{2}}{(a-b)^{2}}$.
Fringe visibility, $V=\frac{I_{\max }-f_{\min }}{l_{\max }+I_{\min }}$
- When entire apparatus of Young's double slit experiment is immersed in a medium of refractive index $\mu$, then 立inge width becomes

$$
\boldsymbol{\beta}^{\prime}=\frac{\hat{\lambda}^{\prime} D}{d}=\frac{\lambda D}{\mu d}=\frac{\beta}{\mu}
$$

- When a thin transparent plate of thichess $t$ and iefractive index $\mu$ is placed in the path of see of the interfering waves, fringe width remains unaffected but the entire pattem shifts by

$$
\Delta x=(\mu-1) t \frac{D}{d}=(\mu-1) t \frac{A}{\lambda}
$$

This shifting is towards the side in which transparent plate is introduced.

## Coherent sources

- The sources of light, which emit continuous light waves of the same wavelength, same frequency and in same phase or having a constant phase difierence are known as coherent sources.
- Two independent sources or light cannot be coherent.


## 

The interference pattern of two identical slits separated by a distance $d=0.25 \mathrm{~mm}$ is observed on a screan at a distance of 1 m from the plane of the slits. The slits are illuminated by manochiomatic light of waveleagth 589.3 nm (scdium travelling perpendicular to the plane of the slits. Bright bands are observed on each side of the central maxima. Calculate the separation between adjacent bright bands?
Solm.: Fringe width $(B)=\frac{D \lambda}{d}=\frac{(\mathrm{im})\left(589.3 \times 10^{-9} \mathrm{~m}\right)}{\left(0.25 \times 10^{-3}\right) \mathrm{m}}$

$$
=2 \times 1{ }^{-3} \mathrm{~m}=2 \mathrm{~mm}
$$

## 

In a 3 DSE, the slits are 2 mm apart and are illuminated with a mixture of two wavelengths $\lambda=750 \mathrm{~nm}$ and $\lambda^{\prime}=900 \mathrm{~nm}$. At what distance from the common central bright finge on a screen 2 m from the slits will a bright finge from one interference pattern coincide with a bright finge from the other?
Solan: The $n^{\text {th }}$ bright fringe of the $\lambda$ paterm and the $n^{\text {tin }}$ bright fringe of the $\lambda^{\prime}$ pattern are situate at

$$
y_{n}=n \frac{D \lambda}{d^{2}} \text { and } y_{n}^{\prime}=n^{\prime} \frac{D \lambda^{\prime}}{a^{\prime}}
$$

As this coincide, $y_{n}=y_{n}^{\prime}$

$$
\begin{aligned}
& \Rightarrow \frac{n D \lambda_{n}}{d}=\frac{n^{\prime}-\lambda_{1}^{\prime}}{2} \\
& \Rightarrow \frac{n}{n^{\prime}}=\frac{\lambda^{\prime}}{\lambda}=\frac{900}{750}=\frac{6}{5}
\end{aligned}
$$

hence the first position where overlapping occur is

$$
y_{5}^{2}=y_{6}=\frac{n D \lambda}{d}=\frac{6(2 \mathrm{~m})\left(750 \times 10^{-9} \mathrm{~m}\right)}{\left(2 \times 10^{-3} \mathrm{~m}\right)}=4.5 \mathrm{~mm}
$$

## 

The intensity of the light coming fom one of the slits in a TISSE is double the intensicy from the other slit. Find the ratio of maxinum intensity to minimum intensity in the interference finge pattern observed.

Soln.: $\frac{I_{\text {max }}}{I_{\operatorname{mans}}}=\frac{\left(\sqrt{I_{1}}+\sqrt{I_{2}}\right)^{2}}{\left(\sqrt{I_{1}}-\sqrt{I_{2}}\right)^{2}}$
Now, $I_{1}=2 I_{2} \Rightarrow \frac{I_{\text {max }}}{I_{\text {mina }}}=\left(\frac{\sqrt{2 I_{2}}+\sqrt{I_{2}}}{\sqrt{2 I_{2}}-\sqrt{I_{2}}}\right)^{2}=\left(\frac{\sqrt{2}+1}{\sqrt{2}-1}\right)^{2}=34$.

## DFFRACTION OF MGNT

* It is the phenomeanon of bending of light arourd cormers of an obstacle or aperture in the path of light. On account of this bending, light penetrates into the geometrical shadow of the costacle. The deviation is more effective when the dimensions of the obstacle or aperture are
comparable to the wavelength of light. This is, therefore, the essential condition for diffraction of light.
- Diffraction of light is of two types :

Fresnel diffraction
Fraunhofer diffraction

- Fresnel diffraction : In tbis case, the source or the screen or both are at finite distances from the aperture or obstacle causing diffraction.
- Fraunhofer diffraction : In tbis case, the source and the screen on which the pattern is observed are at infinite distances from the aperture or the obstacle causing diffraction.


## Diffraction due to a Single Slit

- The diffraction pattern produced by a single slit of width a consists of a central maximum bright band with alternating bright and dark bands of decreasing intensity on both sides of the central maximum.
- Condition for $n^{\text {th }}$ secondary maximum is

Path difference $=a \sin \theta_{n}=(2 n+1) \frac{\lambda}{2}$
where $n=1,2,3, \ldots \ldots$.

- Condition for $\pi^{\text {th }}$ secondary minimum is

Path difference $=a \sin \theta_{n}=n \lambda$
where $n=1,2,3, \ldots \ldots$.

- Width of secondary maxima or minima
where $\beta=\frac{\lambda D}{a}=\frac{\lambda f}{a}$
$a=$ width of slit
$D=$ distance of screen from the slit
$f=$ focal length of lens for diffracted light
Width of central maximum $=\frac{2 \lambda D}{a}=\frac{2 f \lambda}{a}$
- Angular fringe width of central maximum $=\frac{2 \lambda}{a}$.
- Angular fringe width of secondary maxima or minima $=\frac{\lambda}{a}$


## Illustration 19

Light of wavelength 580 mrn is incident on a slit of width 0.30 mm . The observing screen is placed 2.0 m from the slit. Find the positions of the first dark fringes and the width of the central bright fringe.
Soln.: For first dark fringe,

$$
\sin \theta= \pm \frac{\lambda}{a}= \pm \frac{5.8 \times 10^{-7} \mathrm{~m}}{0.3 \times 10^{-3} \mathrm{~m}}= \pm 1.9 \times 10^{-3}
$$

The positions of the first minima, measured from the central axis are
$y_{1}= \pm D \tan \theta \approx \pm D \sin \theta$
As $\theta$ is small, $\tan \theta=\sin \theta$
$\Rightarrow y_{1}= \pm D\left(\frac{\lambda}{a}\right)= \pm 2\left(1.9 \times 10^{-3} \mathrm{~m}\right)= \pm 3.8 \times 10^{-3} \mathrm{~m}$

The positive and negative signs correspond to the first dark fringes on either side of the central bright fringe. Hence, the width of the cen ral bright fringe is given by $2\left\lfloor v_{1} \mid=7.6 \times 10^{-3} \mathrm{~m}=7.6 \mathrm{~mm}\right.$.
Note that this value is much greater than the width of the slit ( 0.3 mm )

## Illustration 20

In a single slit diffraction experiment, first minimum for red light ( 660 nm ) coincides with the first maximum of another wavelength $\lambda^{\prime}$. Then find the value of $\lambda^{\prime}$.
Soln.: For minima, $d \sin \theta=n \lambda$
Here $n=1, \boldsymbol{d} \cdot\left(\frac{y_{1}}{D}\right)=1 \cdot(6600 \AA)$
$y_{1}=\frac{D}{d} \cdot(6600 \AA)$
Now, first maximum is approximately between the first minima and second minima.
$y_{I}=\left(\frac{y_{1}+y_{2}}{2}\right)=\left(\frac{1+2}{2}\right) \cdot \frac{D \lambda^{\prime}}{d}$
As $y_{1}=y_{\mathrm{I}} \Rightarrow \frac{D}{d}(6600 \AA)=\left(\frac{3}{2}\right) \frac{D}{d} \lambda^{\prime}$
$\lambda^{\prime}=\frac{2 \times 6600 \AA}{3}=4400 \AA$

## Fresnel distance

- It is the minimum distance a beam of light has to travel before its deviation from straight line path becomes significant.

$$
\text { Fresnel distance, } Z_{F}=\frac{a^{2}}{\lambda}
$$

## Resolving power

- It is the ability of an optical instrument to produce distinctly separate images of two close objects i.e. it is the ability of the instrument to resolve or to see as separate, the images of two close objects.


## Limit of resolution

- The minimum distance between two objects wbich can just be seen as separated by the optical instrument is known as the limit of resolution of the instrument. Smaller the limit of resolution of the optical instrument, greater is its resolving power and vice-versa.


## Rayleigh's criterion of limiting resolution

- According to Rayleigh, two nearby images are said to be resolved if the position of the central maximum of one coincides with the first secondary minimum of the other and vice versa.


## Resolving power of a microscope

- It is defined as the reciprocal of the minimum distance $d$ between two point objects, which can just be seen through the microscope as separate.

$$
\text { Resclvingpower }=\frac{1}{d}=\frac{2 \mu \sin \theta}{\lambda}
$$

where $\mu$ is refractive index of the medium between object and objective lens, $\theta$ is half the angle of cone of light from the point object, af represents limit of resolation of microscope and $\mu \sin \theta$ is known as the numerical aperture.

## Resolving power of a telescope

- It is defined as reciprocal of the smallest angular separation (d) between two distant objects, whose images are just seen in the telescope as separate.

$$
\text { Resolving power }=\frac{1}{d \theta}=\frac{D}{122 \lambda}
$$

where $D$ is diameter or aperture of the objective lens of the ielescope, $d \theta$ represents limit of resolation of telescope.

## POLARILATHO

- The phenomenon of restricting the vibrations of light (electric vector) in a particular direction, perpendicular so direction of wave motion is known as polarization of light.
* The plane in which vibrations of polarized light are confued is lmown as plane of vibration.
- A plane which is perpendicular to the plane of vibration is bnown as plane of polarization.


## Angle of Polarization

- The angle of incidence for which an ordinary light is completely polarized in the plane of incidence when it gets reflected from a ransparent medium.


## Lates of Malus

- According to law of Malus, when a beam of completely plane polarized light is incident on an analyser, the resultant intensity of light (i) transmitted from the analyser varies directly as the square of the cosine of the angle (0) between plane of transmission of analyser and poiarizer
i.e. $\quad I \propto \cos ^{2} \theta$


## Breuster's faw

- According to Brewster's law, when unpolarized light is incidenî at polarizing angle $\left(i_{p}\right)$ on an interface separativg a rarer medium from a denser madium of refactive index $\mu$, such that

$$
A=\tan i_{p}
$$

then light is reflected in the rarer metium is completely polarized. The reflected and refracted rays are perpendicular t each other.

